

Hypernucleus physics from view point of heavy ion experiment

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**The 1st Symposium on Intermediate-
energy Heavy Ion Collisions, 4/8-11,
Tsinghua Uni., Beijing, China**

Why hypernuclei?

★ QCD theory development

- Micro-lab. with proton, neutron and hyperons;
- YN & YY interaction (strangeness sector of hadronic EOS)

★ Astrophysics

- Hyperons are important for cosmology, physics of neutron stars...

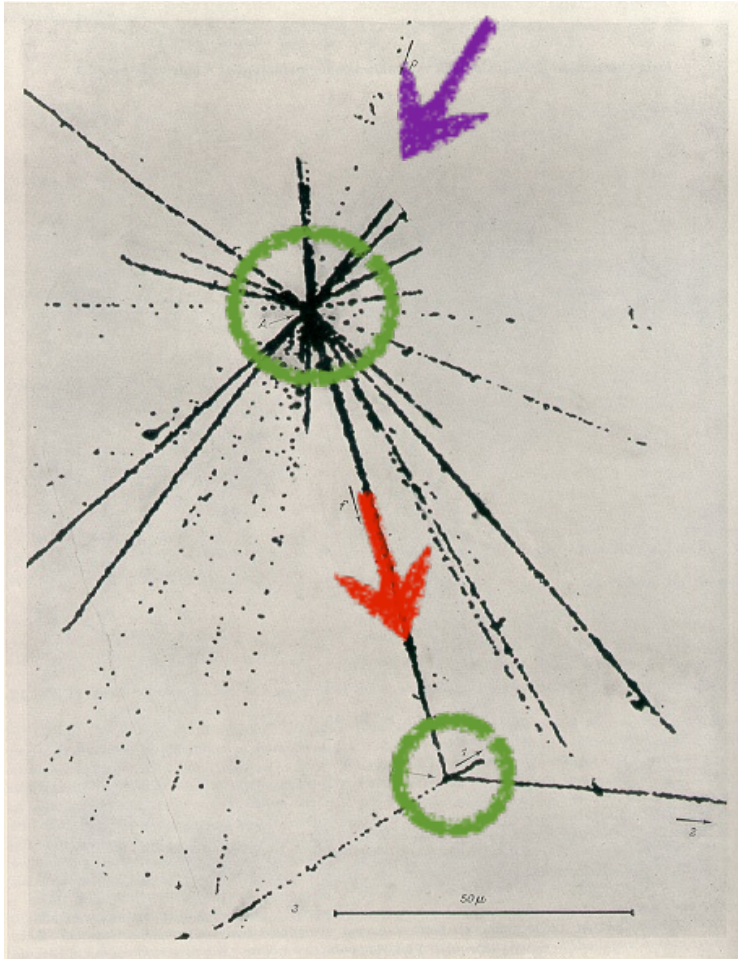
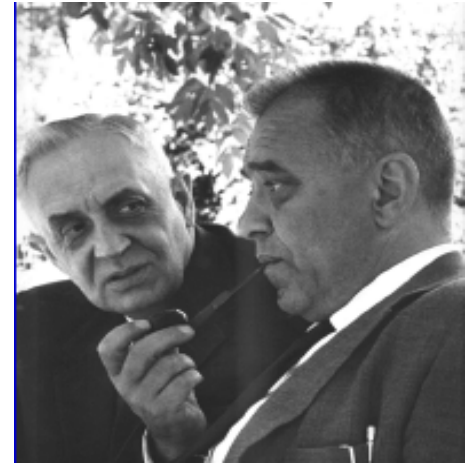
★ Nuclear Physics

- Phenomenology: extension of nuclear charts into strangeness, exotic nuclei...
- Structure theory: nuclear dof for investigating interaction of baryons in nuclei (hyperons – w/o Pauli blocking)
- Reaction theory: new probe for fragmentation of nuclei, phase transition and EOS in hypermatter and finite hypernuclei

The first hypernucleus

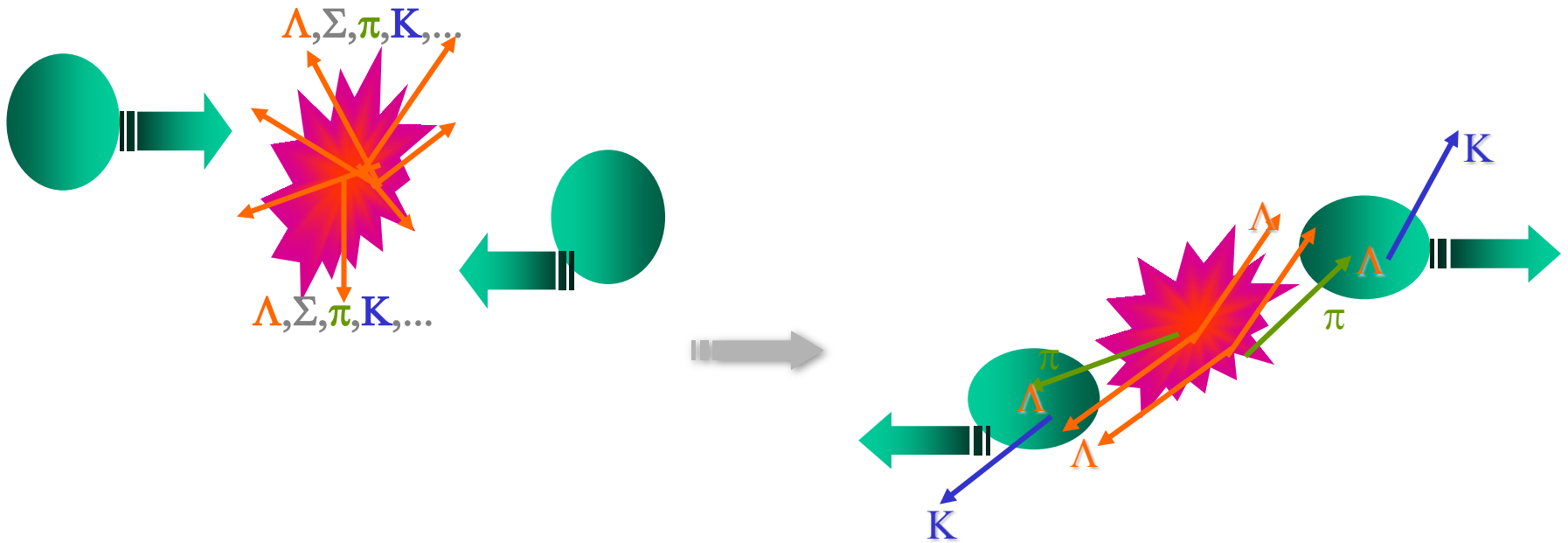
★ The 1st hypernucleus was observed in a stack of photographic emulsions exposed to cosmic rays at about 26 km above the ground.

M. Danysz & J. Pniewski, *Phil. Mag.* **44**, 348 (1953)



- Incoming high energy proton from cosmic ray, colliding with a nucleus of the emulsion, breaks it in several fragments forming a star.
- All nuclear fragments stop in the emulsion after a short path
- From the 1st star, 21 tracks $\rightarrow 9\alpha + 11\text{H} + 1 \Lambda\text{X}$
- The fragment ΛX disintegrates later, makes the bottom star. Time takes $\sim 10^{-12}$ sec (typical for weak decay)
- This particular nuclear fragment, and the others obtained afterwards in similar conditions, were called **hyperfragments or hypernuclei**

Hypernuclei production in HI Collisions



★ Production of hypermatter in relativistic HI and hadron collisions

- Production of strange particles and hyperons by “participants”,
- Rescattering and absorption of hyperons by excited “spectators”

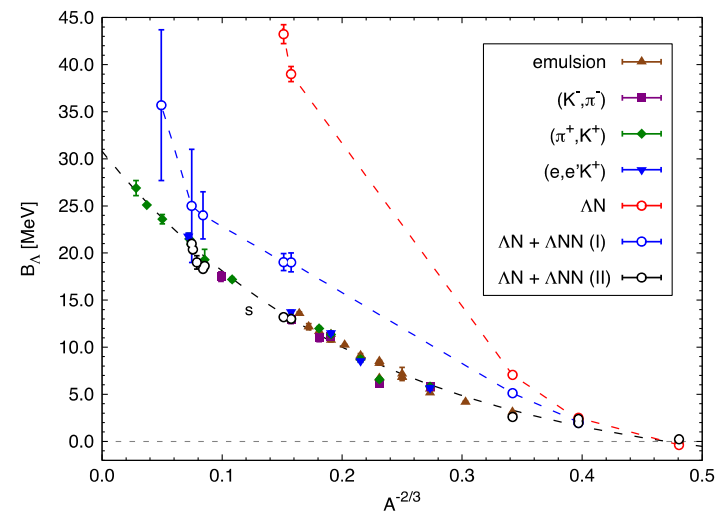
Hyperons in neutron stars

★ “Hyperon puzzle”

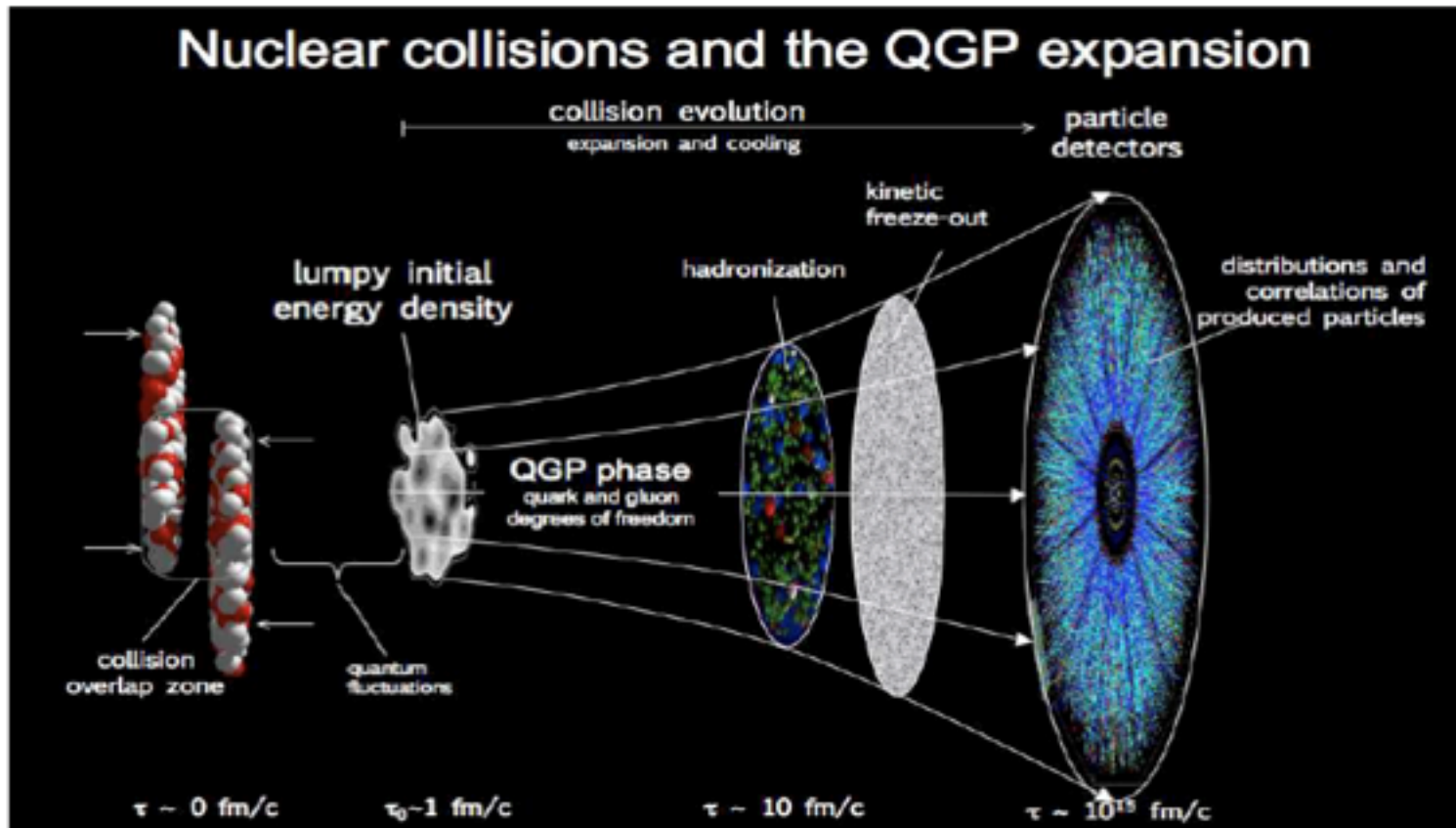
- Hyperons are predicted to exist inside neutron stars at densities exceeding $2-3\rho_0$
- The inner core of NS is so dense, Pauli blocking prevents hyperons from decaying by limiting the phase space available to nucleons
- The presence of hyperon reduces the maximum mass of NS $\sim 0.5-1.2M_0$
- However, new observation for large mass of NS!

P. Demorest et al., Nature 467 (2010) 1081; Antoniadis et al., Science 340 (2013) 448

- ✓ Rijken and Schulze:
Inclusion of YY interactions increase mass of NS
- ✓ Lonardoni et al.,
repulsive ΛNN interactions increase the mass of NS



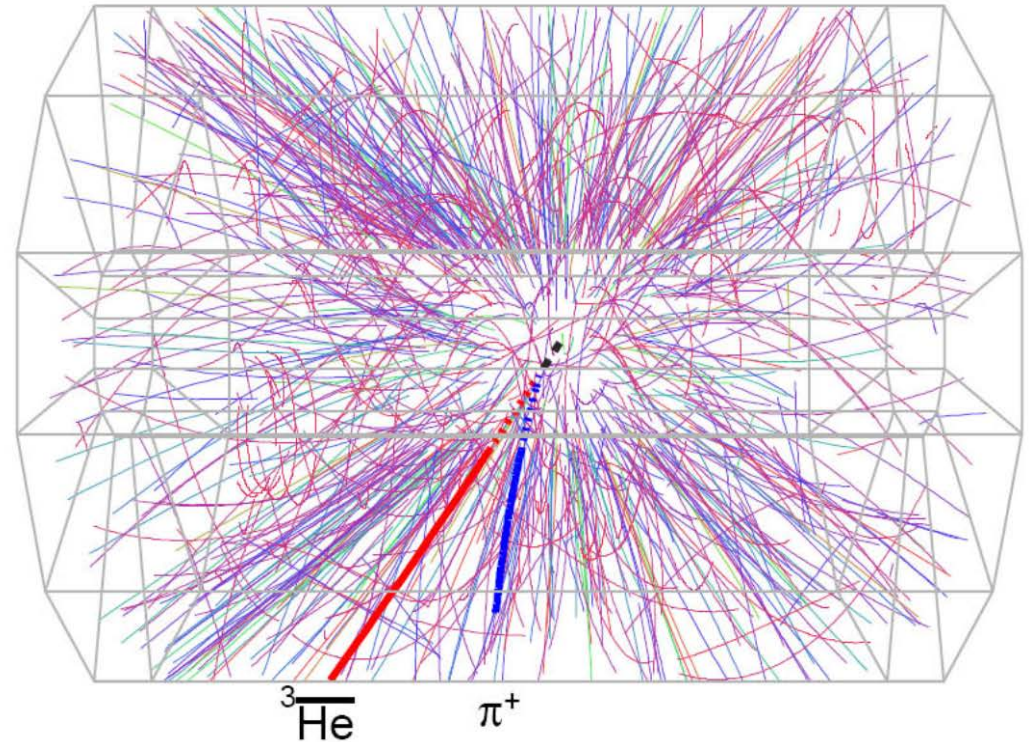
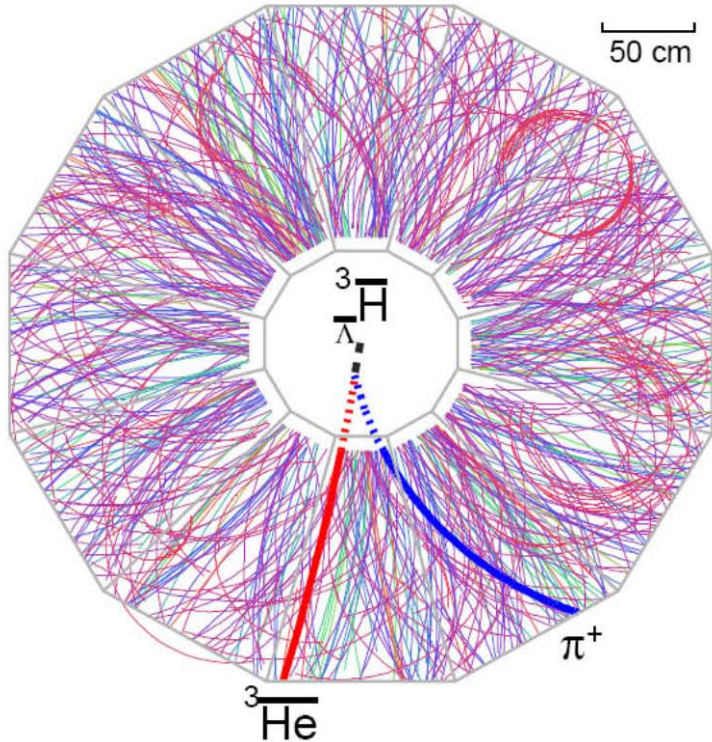
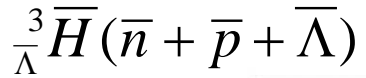
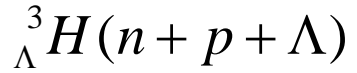
“Stronger constraints on Y-N force are necessary in order to properly assess the role of hyperons in NS”



- ★ Dozens of hyperons production in central heavy-ion collisions at RHIC and LHC
- ★ Excellent secondary vertex reconstruction with STAR and ALICE

Event display

STAR Collaboration, Science 328 (2010) 58



- ★ A beautiful event in the STAR TPC that includes the production and decay of an antihypertriton candidate. (Data taken from Run4 Au+Au 200GeV MB collision)

The mesonic decay of hypertriton

Kamada et al., PRC 57, 1595(1998)

TABLE I. Partial and total mesonic and nonmesonic decay rates and corresponding lifetimes.

Channel	Γ [sec ⁻¹]	$\Gamma / \Gamma_{\Lambda}$	$\tau = \Gamma^{-1}$ [sec]
${}^3\text{He} + \pi^{-}$ and ${}^3\text{H} + \pi^0$	0.146×10^{10}	0.384	0.684×10^{-9}
$d + p + \pi^{-}$ and $d + n + \pi^0$	0.235×10^{10}	0.619	0.425×10^{-9}
$p + p + n + \pi^{-}$ and $p + n + n + \pi^0$	0.368×10^8	0.0097	0.271×10^{-7}
All mesonic channels	0.385×10^{10}	1.01	0.260×10^{-9}
$d + n$	0.67×10^7	0.0018	0.15×10^{-6}
$p + n + n$	0.57×10^8	0.015	0.18×10^{-7}
All nonmesonic channels	0.64×10^8	0.017	0.16×10^{-7}
All channels	0.391×10^{10}	1.03	2.56×10^{-10}
Expt. [6]			$2.64 + 0.92 - 0.54 \times 10^{-10}$
Expt. (averaged) [11]			$2.44 + 0.26 - 0.22 \times 10^{-10}$

$$\frac{1}{\tau} = \sum_i \frac{1}{\tau_i}$$

- ★ Though the mesonic decays are Pauli blocked in heavier hypernuclei, they are the dominant channels in hypertriton.
- ★ In experiment, the 2-body helium3 channel and the 3-body deuteron channel are more easy to access.

Focus on the hypertriton lifetime (1)

- ★ The lifetime measurements are interest especially in view of the short values from early experiments :

The 1st measurements is $(0.95^{+0.19}_{-0.15}) \cdot 10^{-10}$ s from helium bubble chamber, by Block et al., presented in the proceeding of Conference on Hyperfragments at St, Cergue, 1963, p.62

Results from AGS nuclear-emulsion experiments: $(0.9^{+2.2}_{-0.4}) \cdot 10^{-10}$ s,
Phys. Rev.136 (1964) B1803,
from Bevatron and AGS: Phys. Rev.139 (1965) B401
2-body (3 in flight, 4 at rest) $(0.8^{+1.9}_{-0.3}) \cdot 10^{-10}$ s
2-body combined with 3-body (5 in flight, 18 at rest) $(3.4^{+8.2}_{-1.4}) \cdot 10^{-10}$ s

- ★ Nuclear-emulsion with maximum likelihood procedure, Nucl. Phys. B 16 (1970) 46,
 $(1.28^{+0.35}_{-0.26}) \cdot 10^{-10}$ s,

Focus on the hypertriton lifetime (2)

★ But NEW measurements gave different values:

Helium bubble chamber from Argonne ZGS:

$$(2.32^{+0.45}_{-0.34}) * 10^{-10} \text{s}, \text{ PRL } 20(1968)819$$

$$(2.64^{+0.84}_{-0.52}) * 10^{-10} \text{s}, \text{ PRD } 1(1970)66$$

$$(2.46^{+0.62}_{-0.41}) * 10^{-10} \text{s}, \text{ NPB } 67(1973)269$$

Nuclear-emulsion from Bevatron:

$$2\text{-body is } (2.00^{+1.10}_{-0.64}) * 10^{-10} \text{s} \text{ and } 3\text{-body } (3.84^{+2.40}_{-1.32}) * 10^{-10} \text{s},$$

$$\text{and a combined of } (2.74^{+1.10}_{-0.72}) * 10^{-10} \text{s}$$

PRLv20(1968)1383

★ How about the theoretical understanding of these experimental results?

Focus on the hypertriton lifetime (3)

- ★ The hypertriton being a loosely-bound nuclear system, its mean lifetime should not be significantly different from that of the free Lambda

Theoretical calculations from Dalitz et al., initially gave a short value and updated later on a larger value close to the free Lambda's

Phys. Lett. 1 (1962) 58 and Nuovo Cimento A 46,786 (1986)

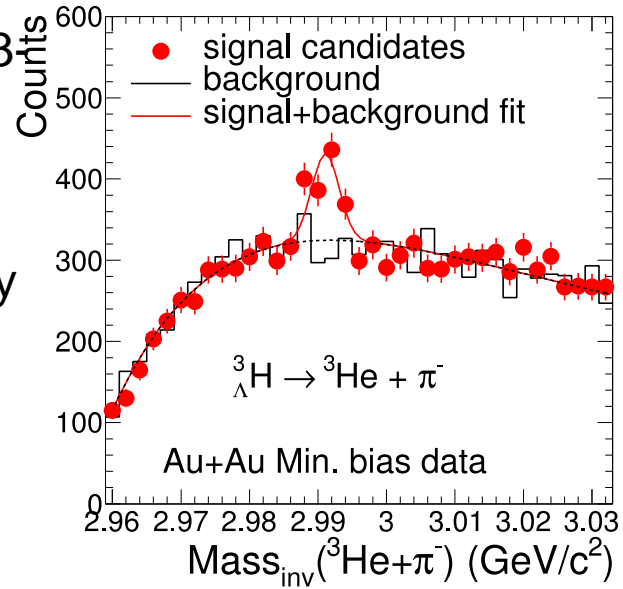
The calculations based on modern 3-body interaction force, the total lifetime is predicted to be $2.56 \cdot 10^{-10}$ s, Phys. Rev. C 57 (1998) 1595

- ★ The hypertriton lifetime data are not sufficiently accurate to distinguish between model, more precise measurements are needed.

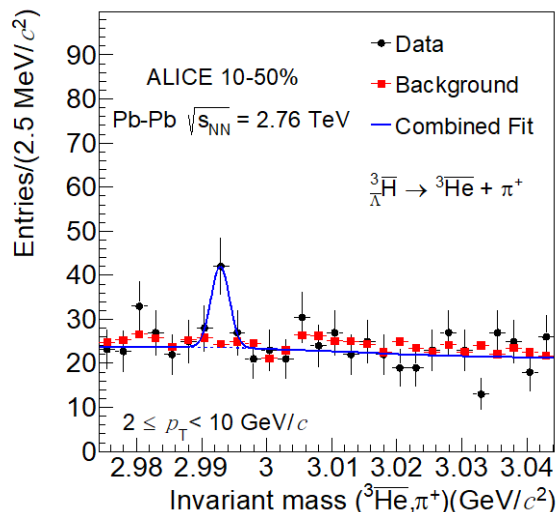
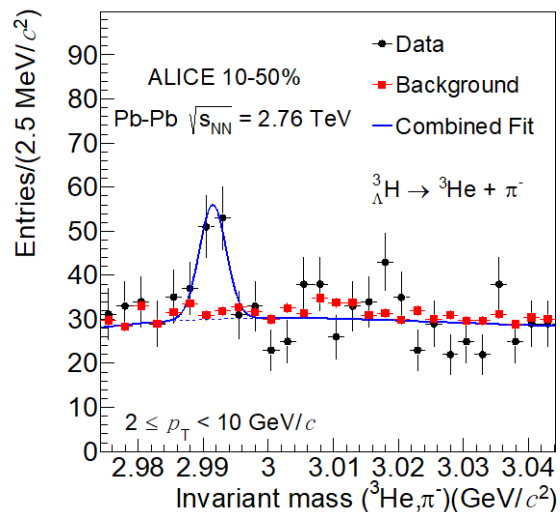
★ Inv. Mass with 2 and 3 body analysis

★ Lifetime is significantly shorter than free lambda

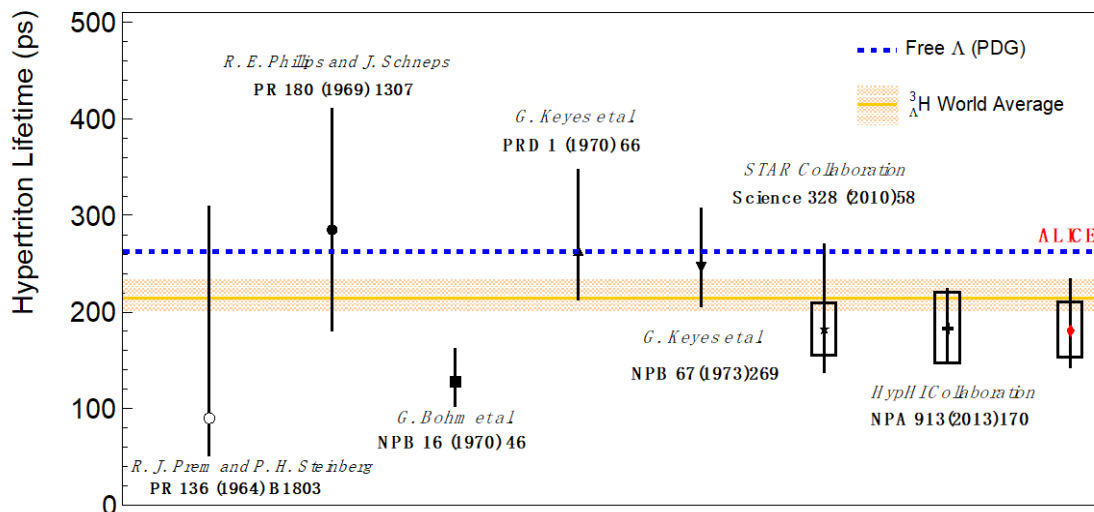
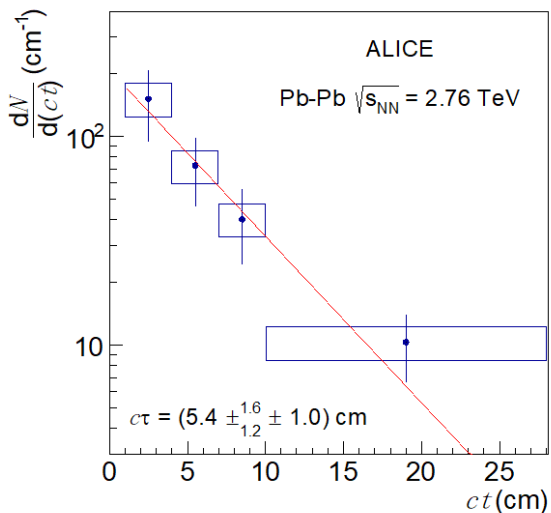
STAR Col. arXiv:1710.00436

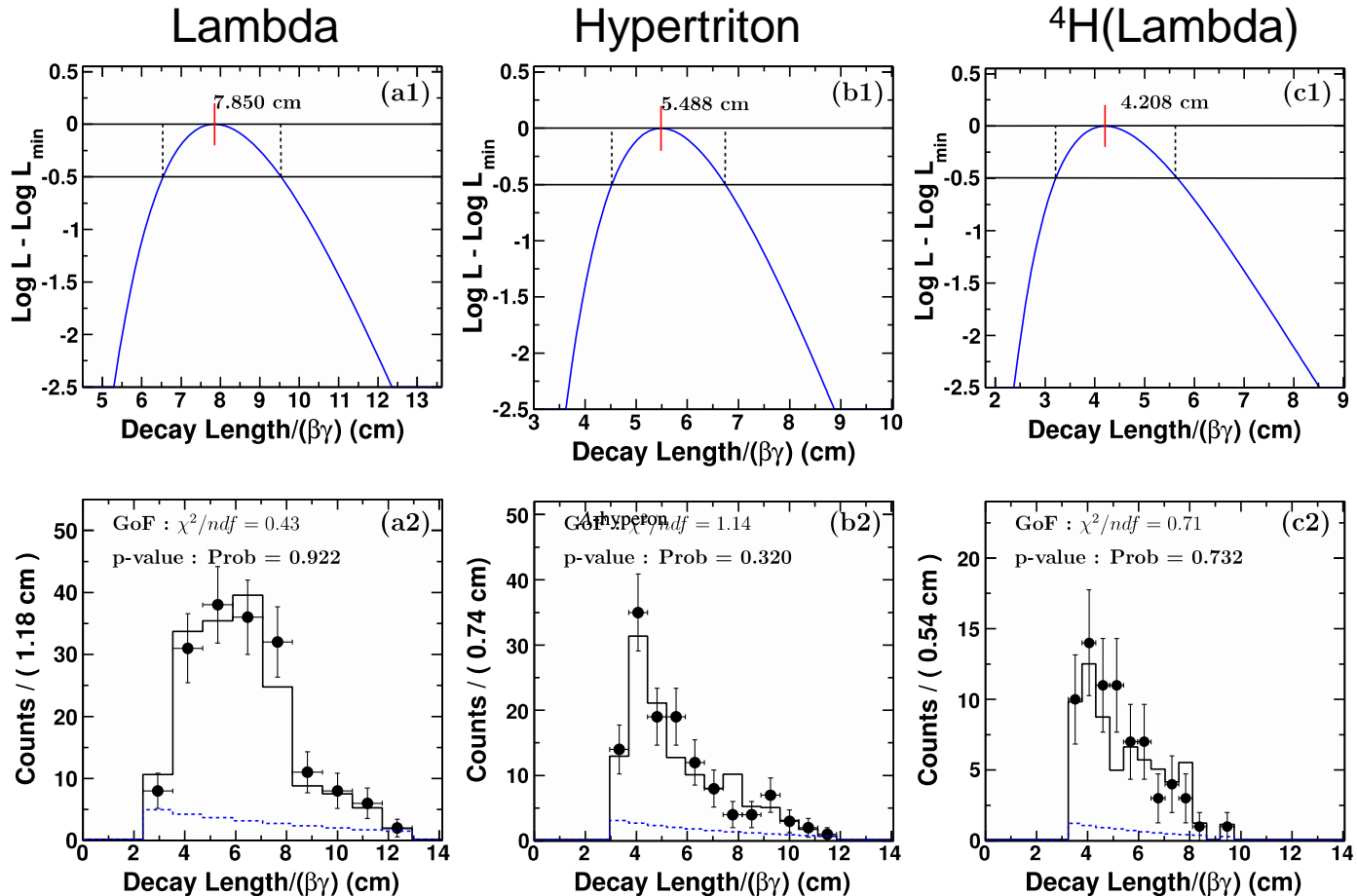


Measurement from LHC-ALICE



ALICE Collaboration
PLB 754, 360 (2016)



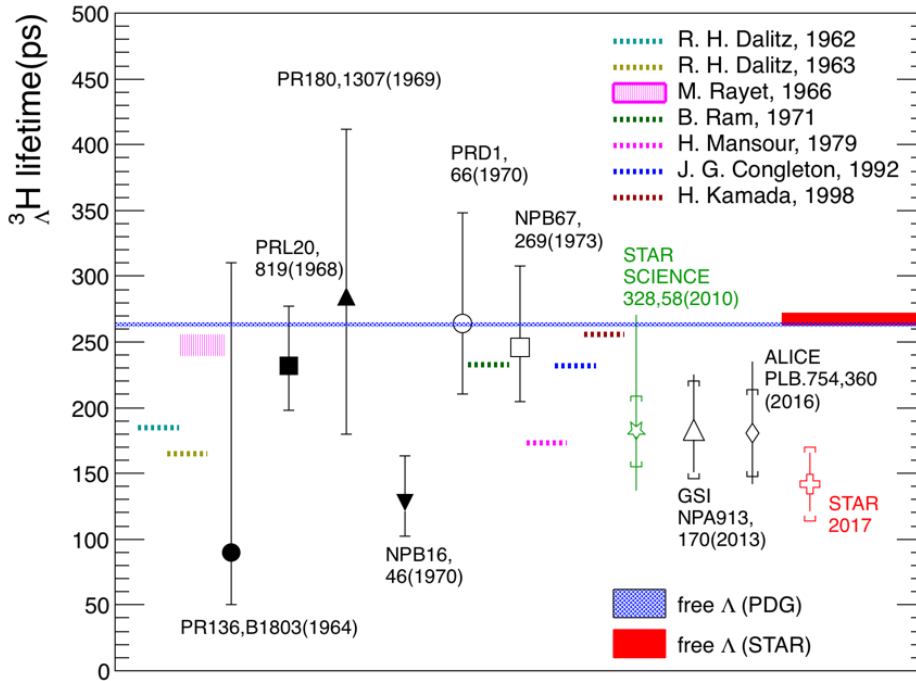


★ Hypernuclear spectroscopy at GSI: ^6Li projectiles on ^{12}C target at 2 A GeV presents the hypertriton lifetime measurement from 2-body channel:

Nucl. Phys. A 913(2013)170

$$\tau = 183 \pm_{32}^{42} \pm 37 \text{ ps}$$

How to understand the data?

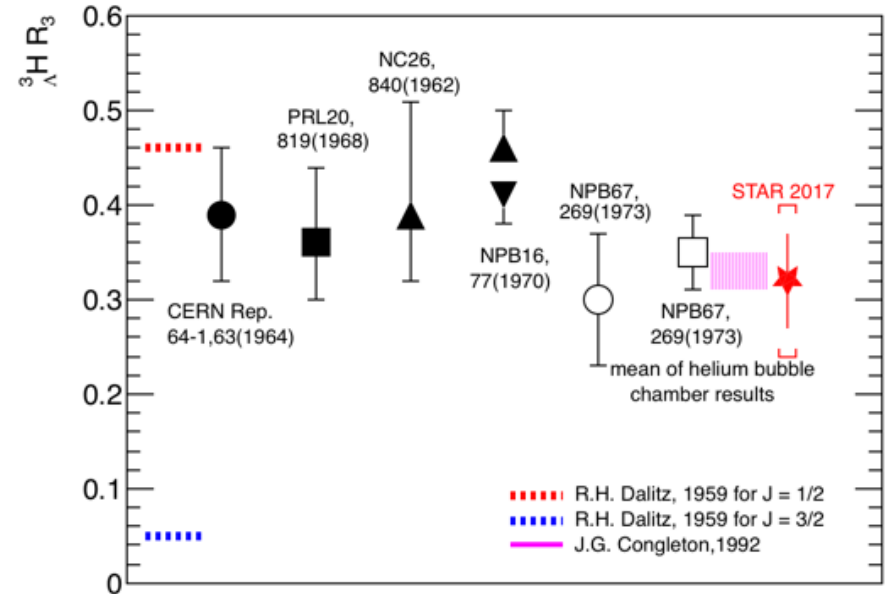


In 2016, Rev. OF. Mod. Phys. 2016 Jul-Sep.

“The discrepancy between the lifetimes measured in HIC and the lifetime prescribed by theory is disturbing, posing a major problem for the understanding of hypertriton, the lightest and hardly bound hypernucleus. More work is necessary to understand the HI lifetime results.”

- Strong Y-N interaction in the hypernucleus system?
- The discrepancy may be in the B_{Λ} ?
- Spin assignment : favor $\frac{1}{2}$

STAR Col. arXiv:1710.00436



We can do more on the B_{Λ}

★ The early data suffers from large statistical uncertainty!

If charge symmetry holds for the ΛN interaction, the ${}^4_{\Lambda}\text{H}$ and ${}^4_{\Lambda}\text{He}$ hypernuclei members of an isotopic spin doublet, should have equal binding energies (with contributions from the distortions of the core nuclei and the Coulomb effect) taken into account. Definite deviations from this prediction indicate that the B_{Λ} value for ${}^4_{\Lambda}\text{He}$ have been reported first by Raymund [14] and confirm

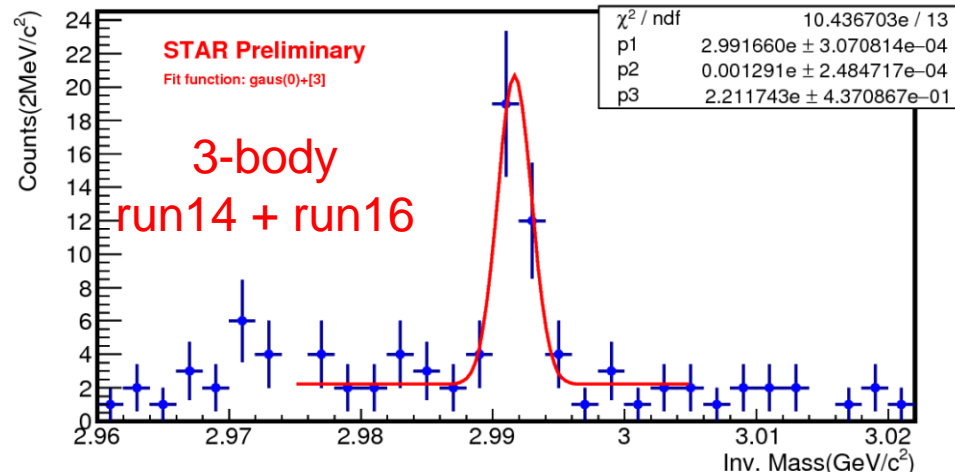
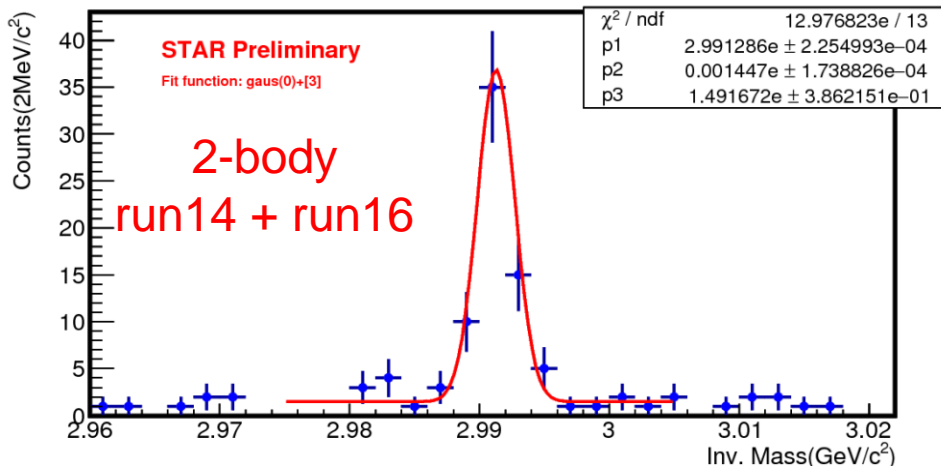
- a) G. Bohm et al., Nucl. Phys. B4 (1968) 511
- b) This work : M. Juric, G. Bohm et al., Nucl. Phys. B52 (1973) 1

for the deuteron mass. The two-body decay events give $B_{\Lambda} = 0.25 \pm 0.31$ MeV, while the combined decays give $B_{\Lambda} = -0.07 \pm 0.27$ MeV. These results should be compared to the two emulsion measurements 0.06 ± 0.06 [20] and 0.24 ± 0.12 MeV. [21]

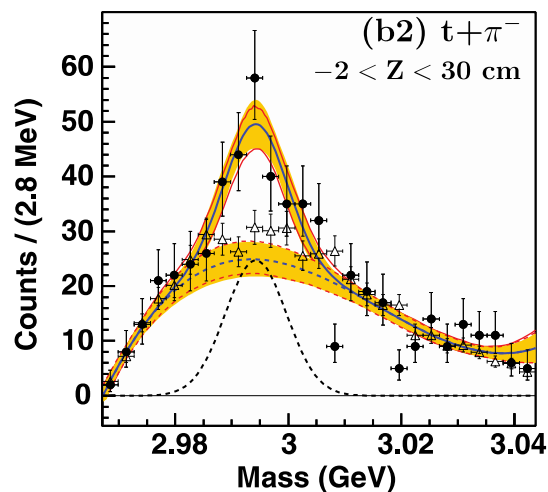
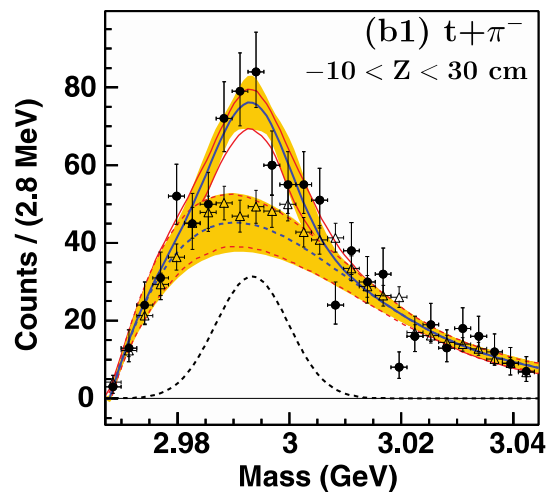
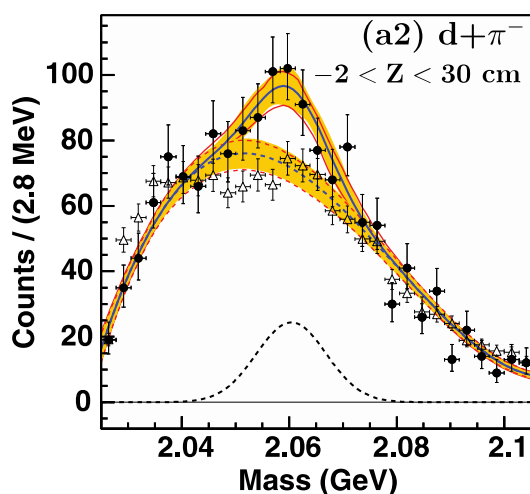
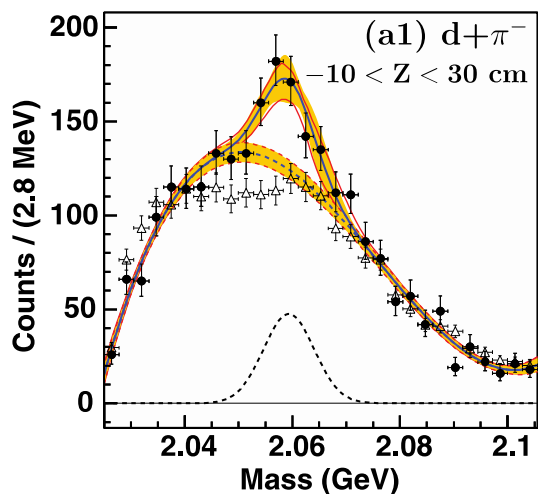
G. Keyes et al., PRD 1 (1970) 66

Dalitz, Nucl. Phys. A 754 (2005) 14c, "I feel that we are far from seeing the end of this road. A good deal of theoretical work on this 3-body system would still be well justified."

★ With micro-vertex detector in STAR, hypertriton is background free. Systematic uncertainty under control!



(nn Λ) signal in HI reaction?



HypHI exp. observed a signal in the invariant mass distribution of $d+\pi^-$ and $t+\pi^-$ channel.

PRC 88,041001R(2013)

Data from HIRES Col.
Excluded the (Lp) candidates

PLB 687, 31(2010)

PRD 84, 032002(2011)

A possible interpretation might be the two- and three-body decays of an unknown bound state of 2 neutrons associated with a Lambda:

$$\text{via } {}^3_{\Lambda}n \rightarrow t + \pi^- \text{ and } {}^3_{\Lambda}n \rightarrow t^* + \pi^- \rightarrow d + n + \pi^-,$$

★ Hypertriton is a local baryon-strangeness correlation system

$$S_3 = \frac{{}^3\Lambda\text{H}}{{}^3\text{He} \times \Lambda / p}$$

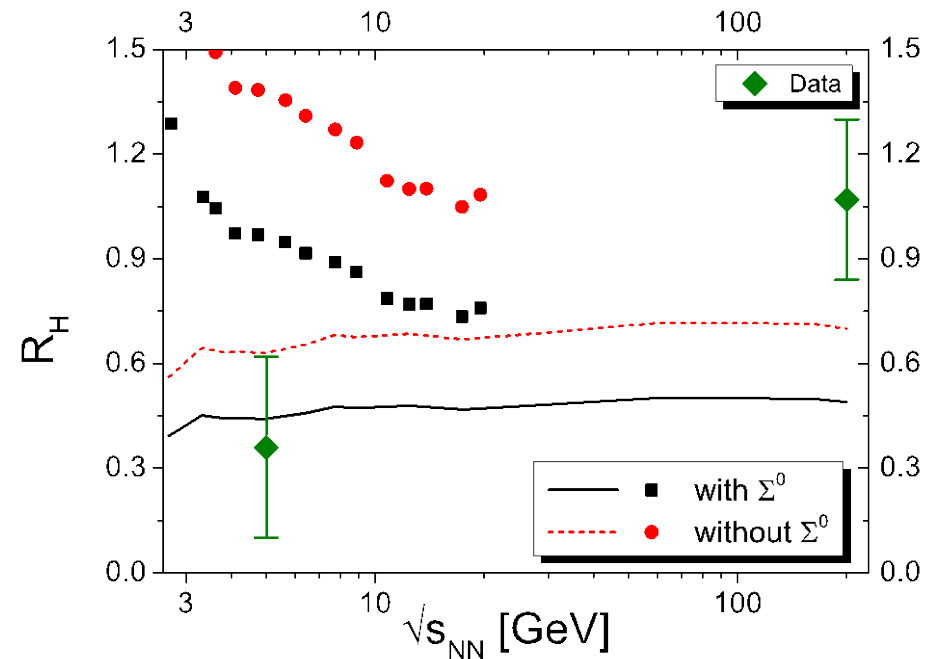
★ Strangeness Population Factor

★ It is predicted that the beam energy dependence of S_3 would behave differently in pure hadron gas and QGP

S.Zhang et al., PLB 684 (2010) 224

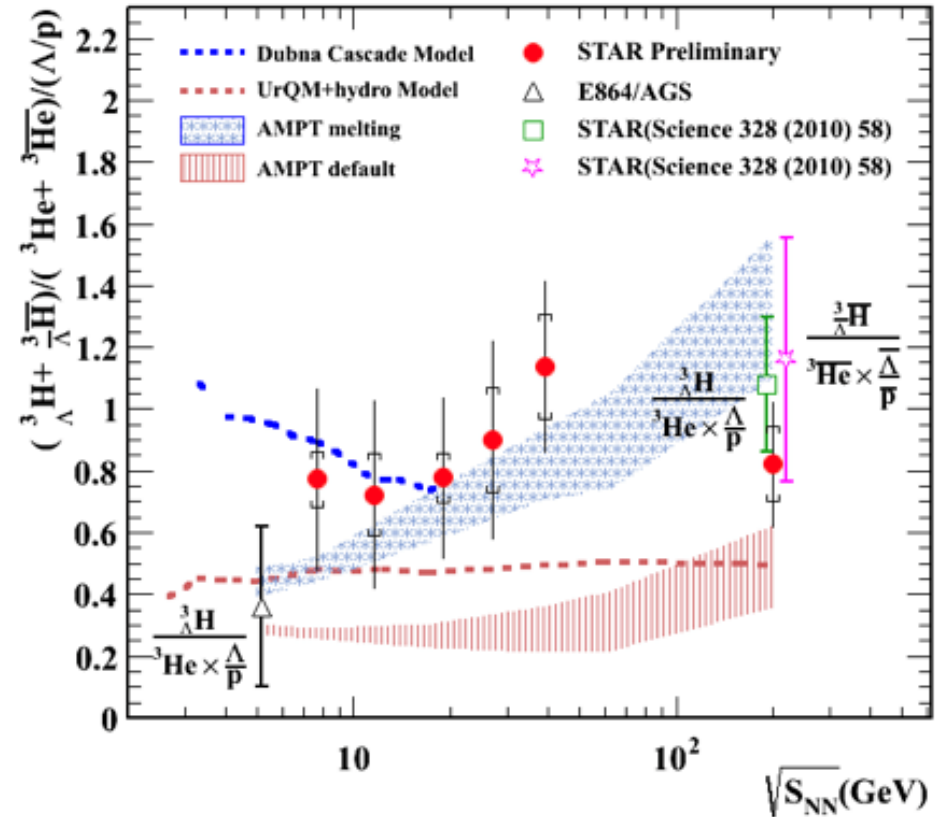
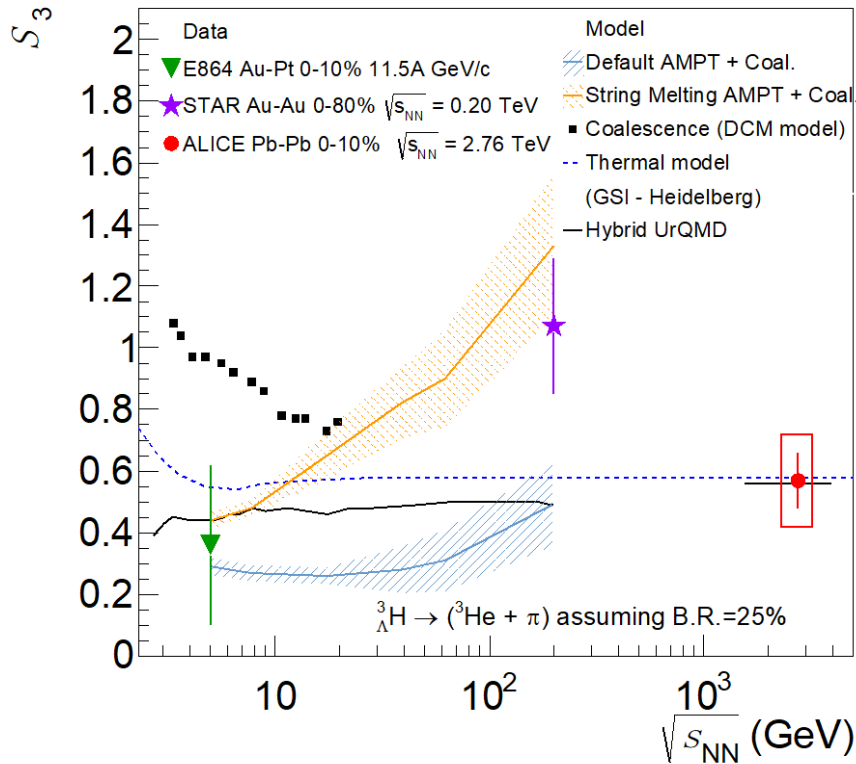
J. Steinheimer et al., PLB 714 (2012) 85

...



Beam energy dependence of S_3

ALICE Collaboration PLB 754, 360 (2016)



★ LHC energy seems to close to thermal model prediction

★ STAR BES data indicates an increasing trend with the increase of energy

Concluding remarks

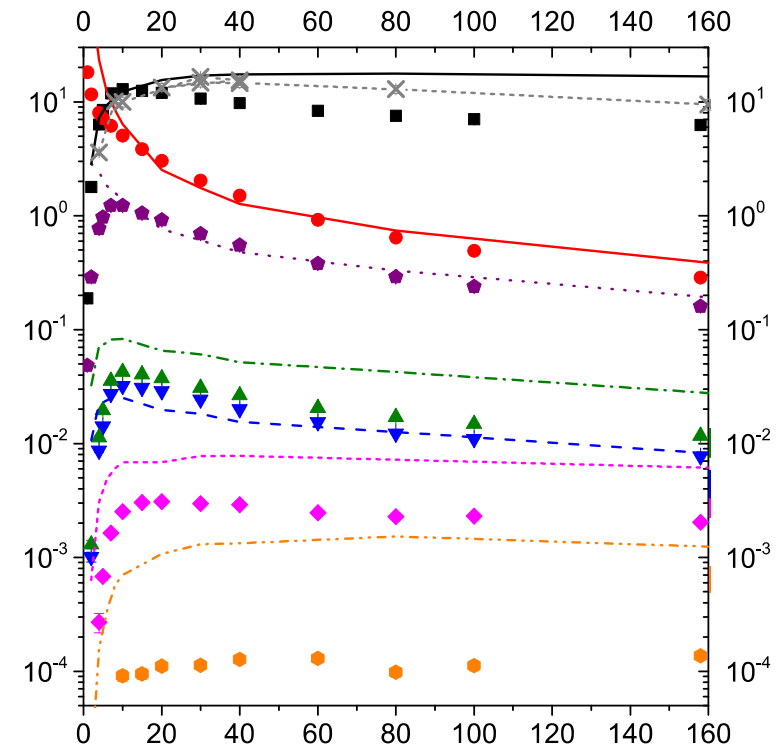
- ★ Hypernuclei is a control tool to study Λ -N interaction in lab.
- ★ Measurements of hypertriton lifetime have been an interesting project in the field. Independent exp. present different results.
- ★ Several theoretical interpretations have been achieved in the field and tend to conclude a value close to the free Λ 's
- ★ New and precise measurements (>600 signals) in heavy-ion collisions have been released. The discrepancy among different exp. is still there, the hypertriton lifetime is still a puzzle. Future direction may be related to the binding energy of light hypernuclei system

Future direction (1)

★ STAR BES-II at 2019 and 2020

- detector upgrades
- low energy electron cooling
- rich hyperon production

Collision Energies	Proposed Event	BES-I Event
7.7	100	4
9.1	160	N/A
11.5	230	12
14.5	300	20
19.6	400	36



J. Steinheimer et al., Phys. Lett. B 714 (2012) 85

Future direction (2)

Proposed (π^- , K^0) reaction on nuclear targets for precise determination of the lifetime of the hydrogen hyperisotopes and other neutron-rich Λ -hypernuclei at J-PARC

M. Agnello et al., NPA 954 (2016) 176

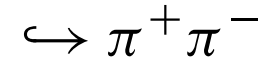
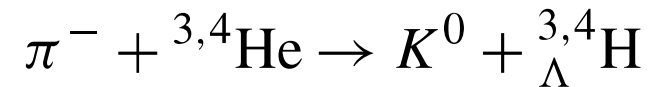
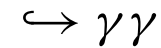
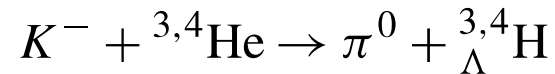
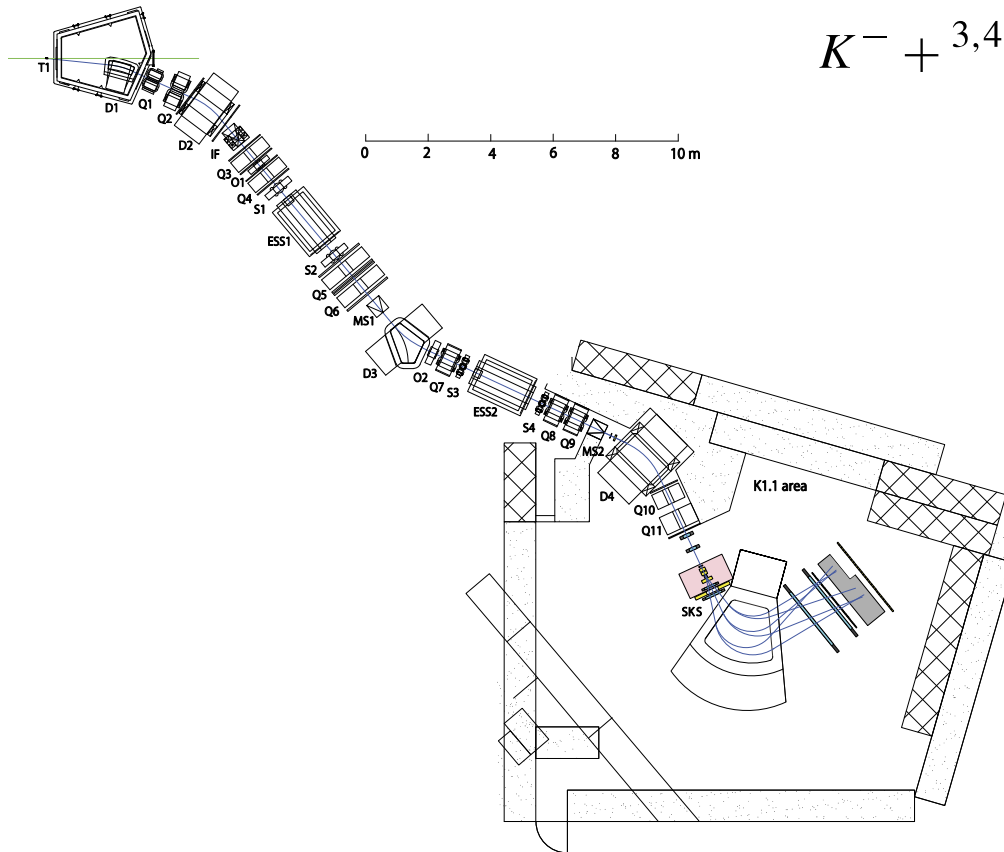
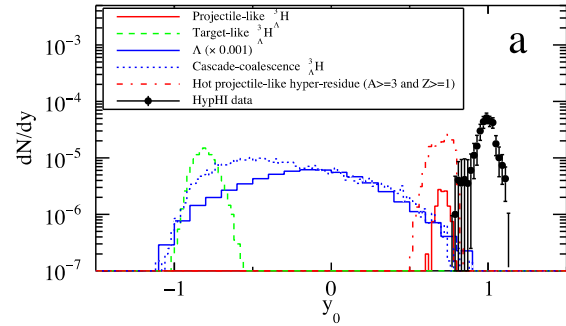


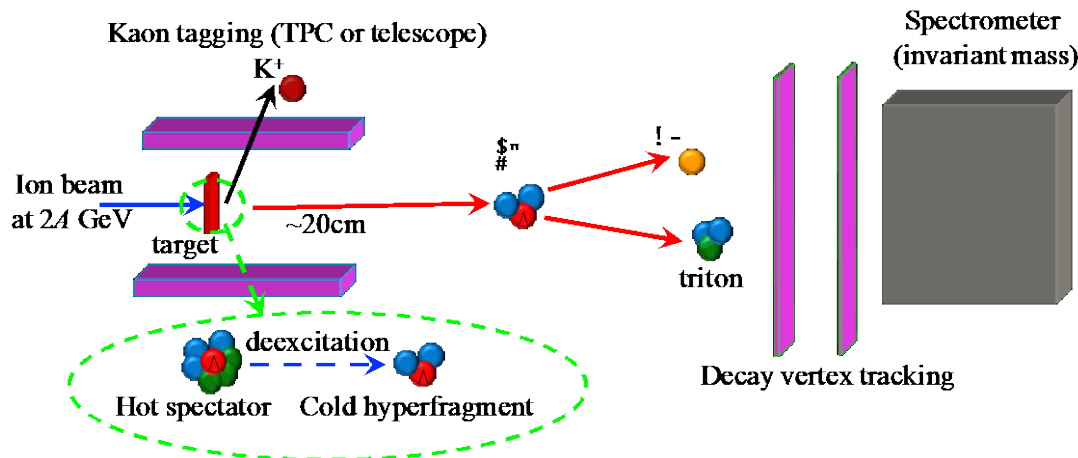
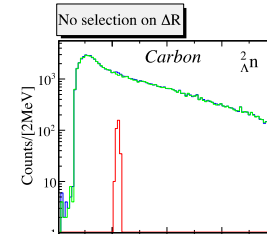
Fig. 4. Layout of the J-PARC K1.1 beam line and K1.1 experimental area. From [46].

Future direction (3)

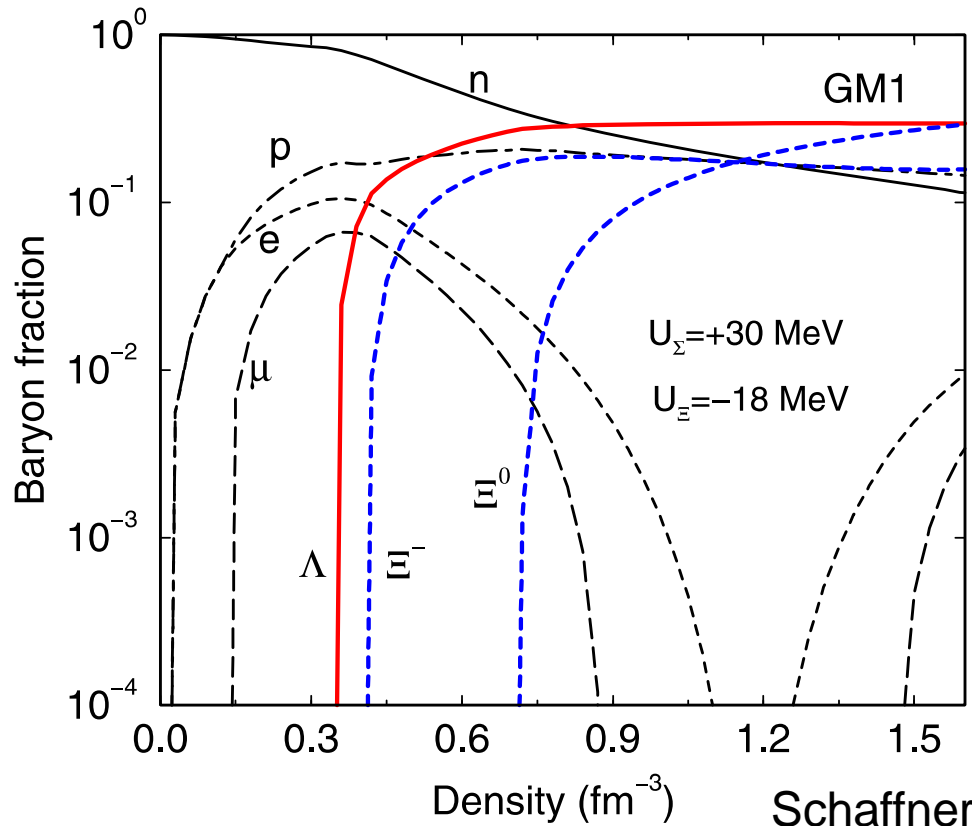


★ Light hypernuclei production in peripheral ion collisions

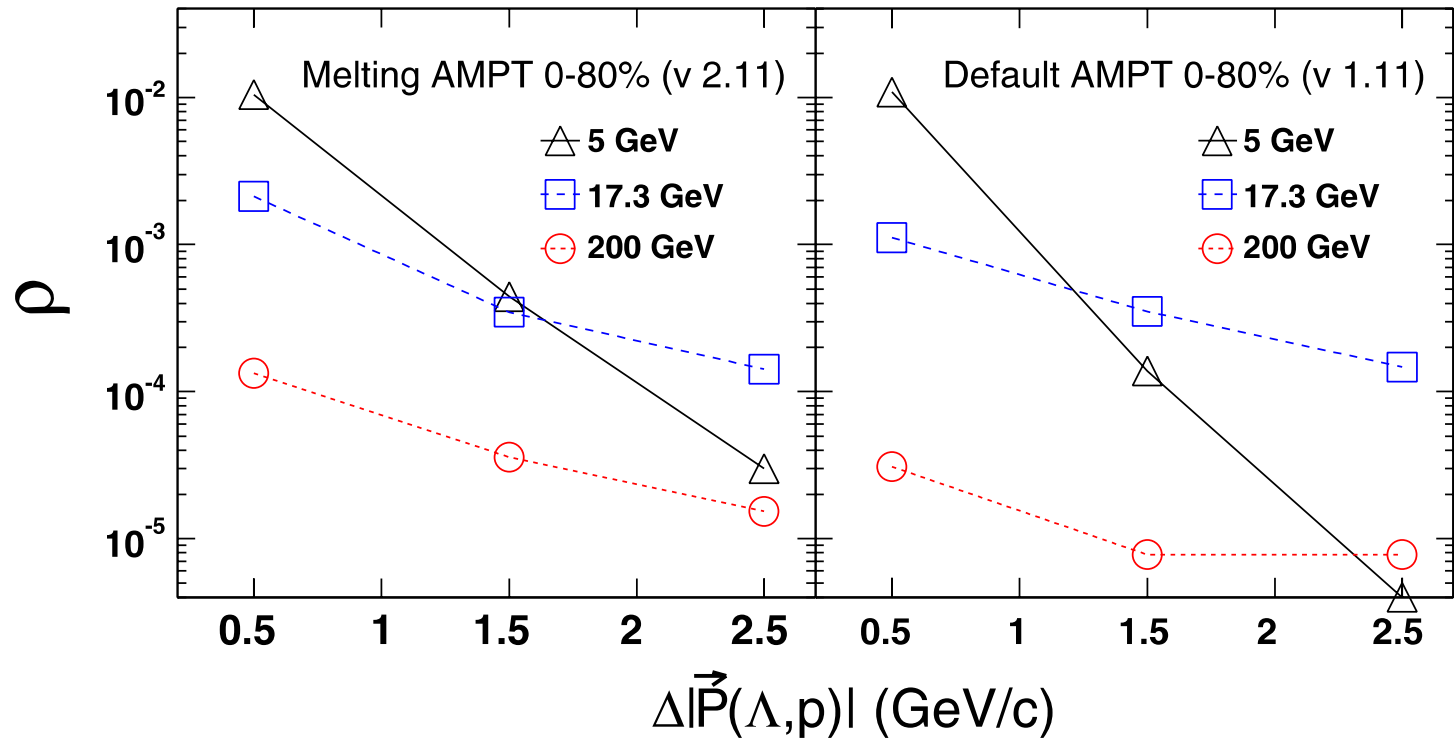
arXiv: 1712.04658



Backup slides

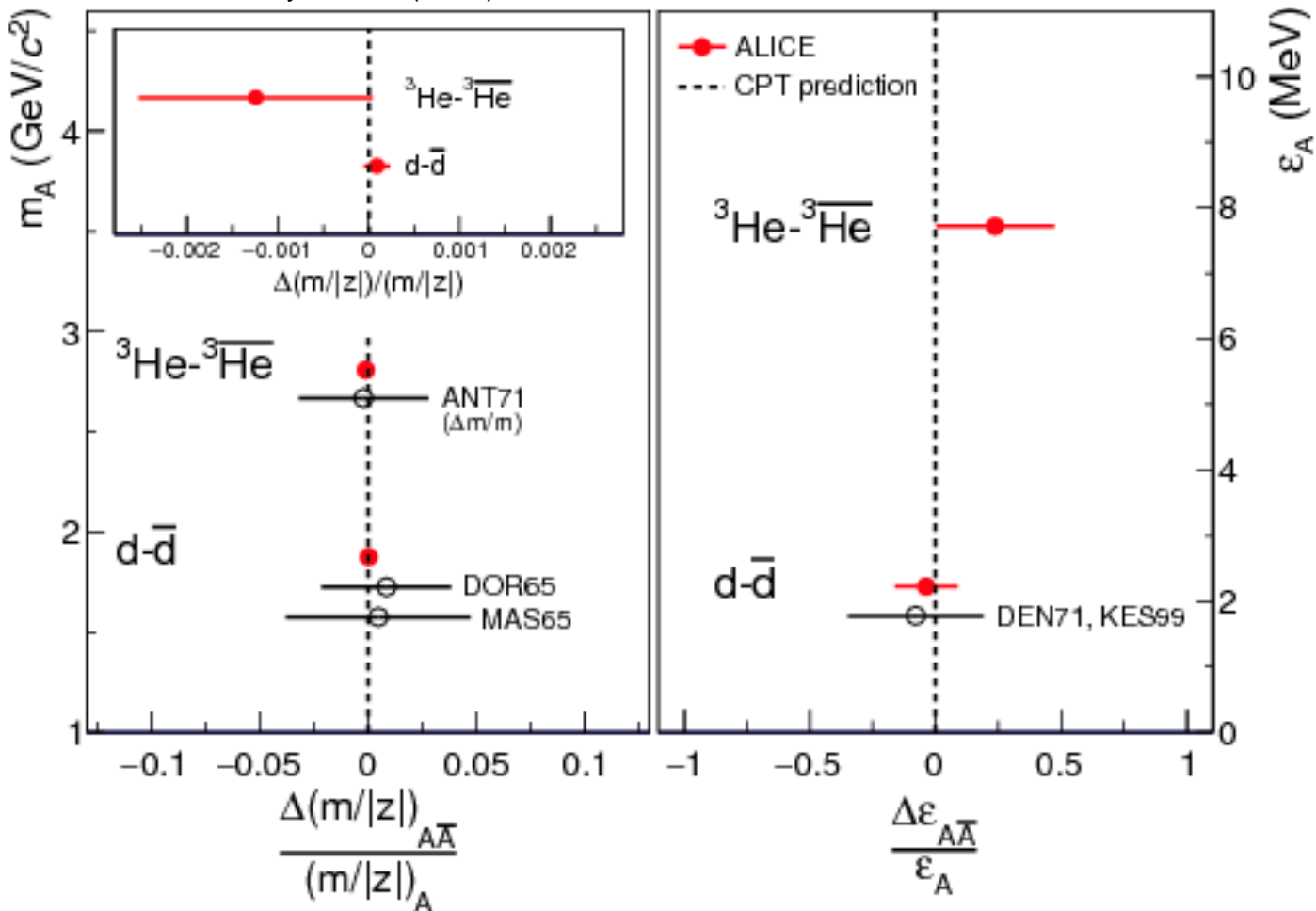


Neutron-star matter fractions of baryons and leptons, calculated as a function of density.



Probing CPT invariant in nuclear force

ALICE Collaboration Nature Physics, 11 (2015) 811



★ Mass-over charge ratio difference confirms CPT invariance to an unprecedented precision in the sector of light nuclei